

FRACTAL S.L.N.E

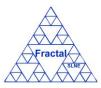


A technological company for astronomical Instrumentation and Software

https://www.fractalslne.es/

Marisa García Vargas, marisa.garcia@fractal-es.com

Introducing FRACTAL

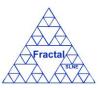


- FRACTAL S.L.N.E. is a private technological company founded in August 2005. (> 17 years)
- The motivation was to keep the know-how accumulated by our team during about 25 years working at ESA, public Research Centers and universities, and the GTC Project Office.
- We focus our services mainly on the Research Centers and Universities with Physics, Engineering and Astronomy Departments, although due to our skills and previous experience we have worked mainly for professional Astrophysics (90%).
- Our best referees are our customers:



http://www.fractal-es.com Contact: Marisa García Vargas marisa.garcia@fractal-es.com

Who are us?: FRACTAL Team



A group of people covering:

- Astronomy
- Management
- System Engineering
- Optics & Adaptive Optics
- Opto-mechanics and Mechanics
- Cryogenics, CCDs and Data Acquisition Systems
- Control Software
- Data Bases
- Telescope Control Systems
- Data Reduction. Pipelines

Full-time, part-time consultants

Most of us worked many years (> 9) at the GTC Project Office in the Instrumentation, Control and System Engineering groups

Also at ESA, ESO and EUMETSAT

Contact: Marisa García Vargas marisa.garcia@fractal-es.com

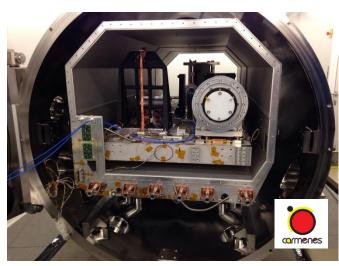


FRACTAL Services



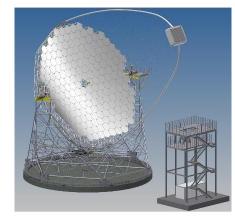
- Management and System Engineering services and products
- Instrumentation: Consultancy, Design and Development
 - Design from scientific requirements definition to conceptual, preliminary and detailed design.
 - Optics (including wide experience on Fibers and Gratings) and Opto-mechanics
 - Cryogenics
 - Integration and tests at laboratory, and commissioning at observatories
- Software development for Scientific Applications
- Scientific operations (Astronomy) and Technical attendance at customer's site
- Specialized Training (general or in-house courses, participation in masters)
- Web services: public pages and customized-Intranet development and hosting
- Partnership with Wasatch Photonics (VPH) and distributor for Europe
- <u>http://www.fractalslne.es/</u> <u>http://www.fractalslne.es/files/FRACTAL-Services-Instrumentation-and-software-general.pdf</u>
- <u>Video with FRACTAL activities: https://youtu.be/161gsgbnRNU</u>

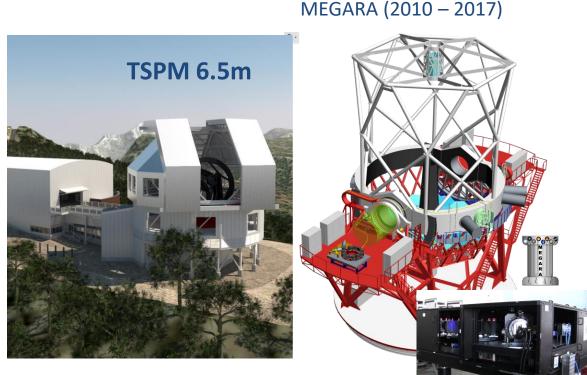
Management and System Engineering services - RAMS plans



Management and System Engineering CARMENES (2013 – 2015)

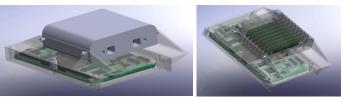
Safety Study for LST-CTA





Management and System Engineering TSPM (IA-UNAM, SO/UA, SAO, INAOE), 2016 - 2020

Management support in the TTT project (SEOSAT, INTA)



Management and System Engineering

LDT Precursor developed for NANOSAT 1B



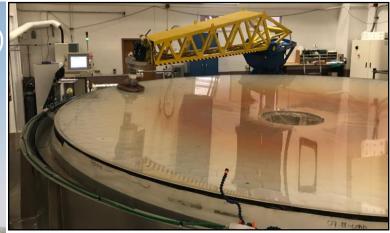


Participation in Large Projects

6.5m Telescopio San Pedro Mártir (Baja California) Detailed design phase. Finished)

- Management
- System Engineering (also SE tools)
- Consultancy in Optics, Mechanics and Civil Work

Partners: IA-UNAM (México) INAOE (México) Steward Observatory/University of Arizona Smithsonian Astronomical Observatory (Harvard, Cambridge Mass)





Calar Alto high-Resolution search for M dwarfs with Exoearths with Near-infrared and optical Echelle Spectrographs

Two fibre-fed stabilised échelle spectrographs (R=82,000) at the coudé thermal rooms of the 3.5 m telescope

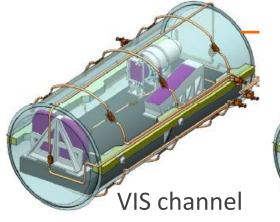
(Detailed Design to Commissioning)

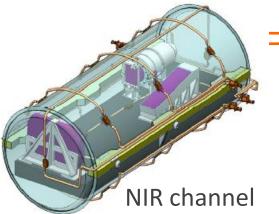
Commissioning: May - December 2015 Regular operations since: January 1st 2016

- Project Manager and System Engineer
- SE tools and web services
- Detector characterization and Data Acquisition system









https://carmenes.caha.es/





CARMENES Detector characterization

spectrógrafo en MEGARA @ GTC TC de esolución para













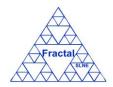


Project Manager

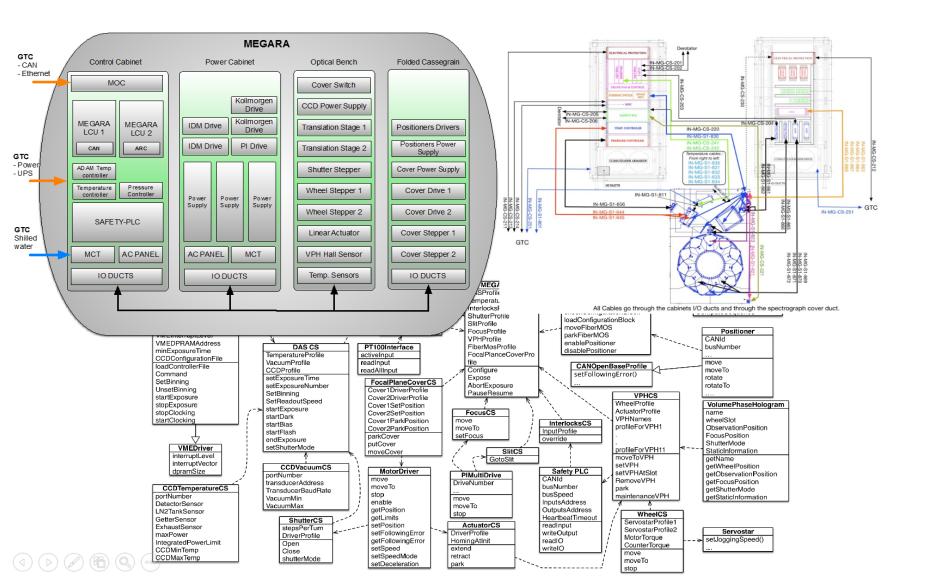
stronomía

- System Engineer (also SE tools and web services)
- Optics, Mechanics, Detector Design and AIV
- Many packages of the Control System
- Detector and Data Acquisition system AIV, tests
- AIV (complete) and commissioning at the Observatory Video on Integration https://www.youtube.com/watch?v=xyBV0b5okwl

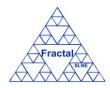
MEGARA Mechanisms control system

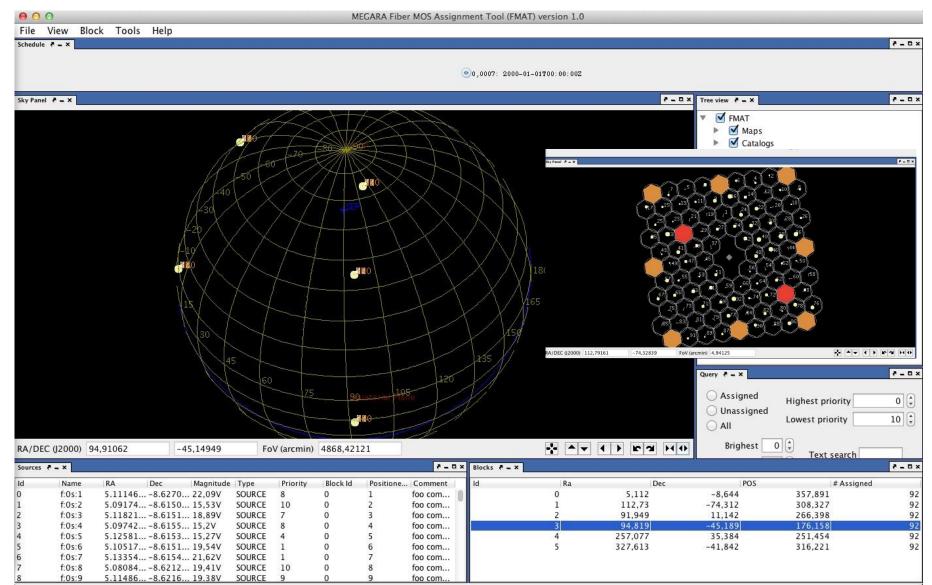


Customer: UCM

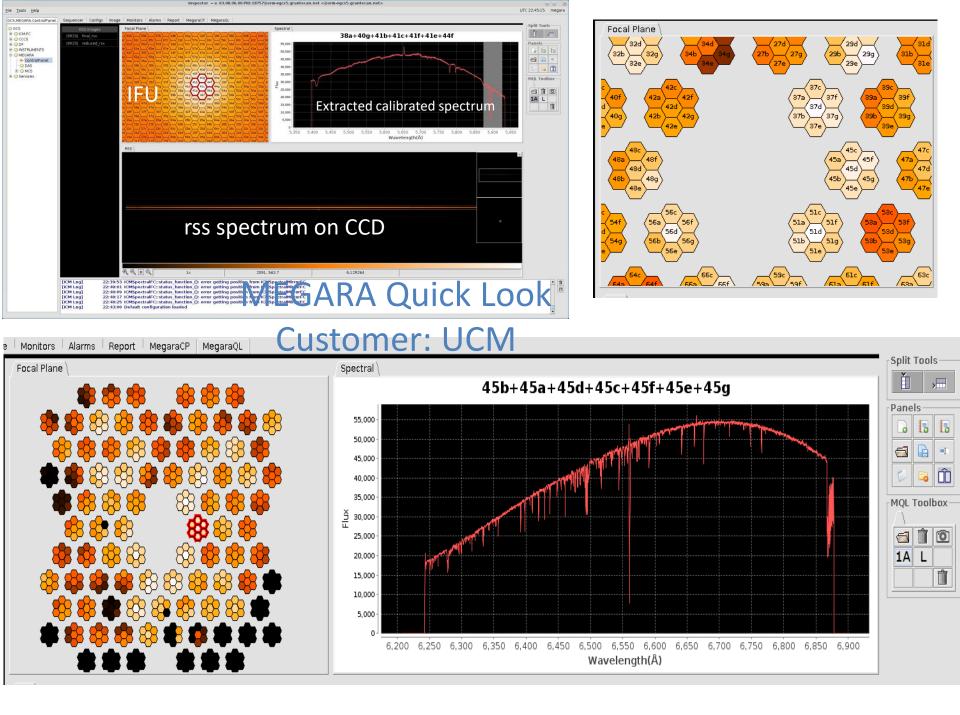


MEGARA Fiber MOS Assignment Tool (FMAT) Customer: IAA / UCM





http://www.fractal-es.com



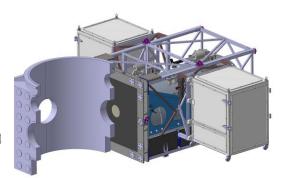
SCORPIO for Gemini Optics/Mechanics/Thermal WPs

Customer: SwRI (San Antonio, TX, USA), currently at subsystems AIV stage

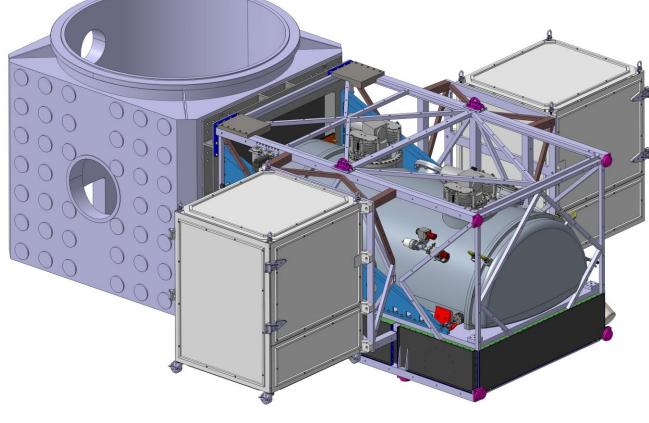
- Management and System engineering of the Optics/Mechanics/Thermal WPs
- Optics design, manufacturing supervision and integration
- Mechanics & Thermal subsystems design, manufacturing supervision and integration
- Manufacturing subcontracts
- Instrument assembly, integration and verification.







FRACTAL Contract: 2017 - 2023



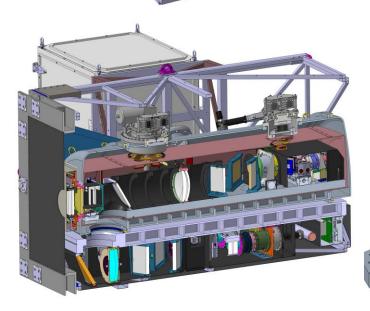
VIS arms: g, r, i, z



THE GEORGE WASHINGTON UNIVERSITY WASHINGTON, DC

STScI

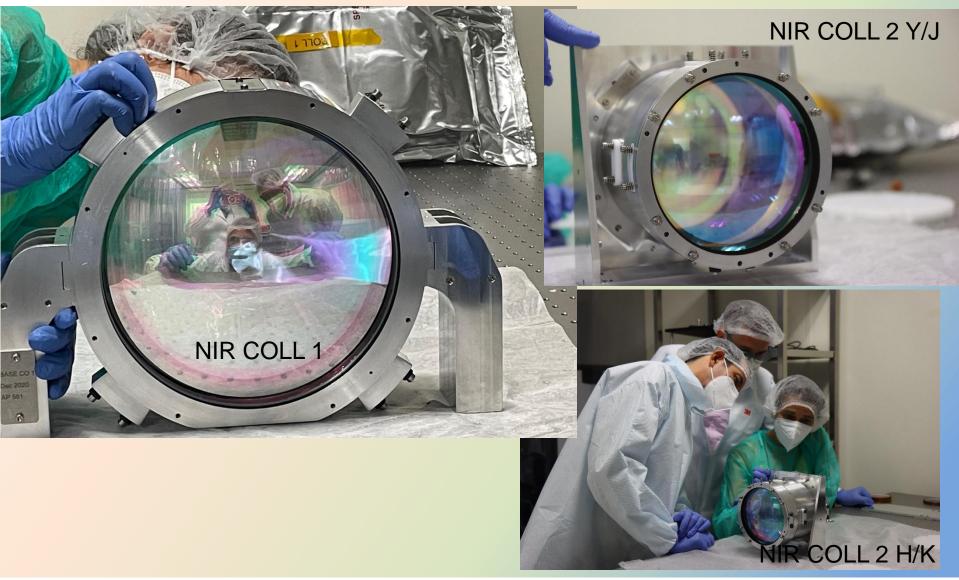
NIR arms: Y, J, H, K





NIR Collimator and camera











NIR beamsplitter YJ/HK @LICA

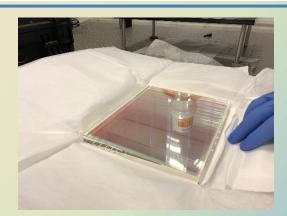


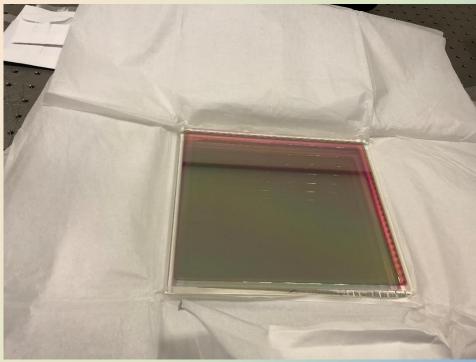












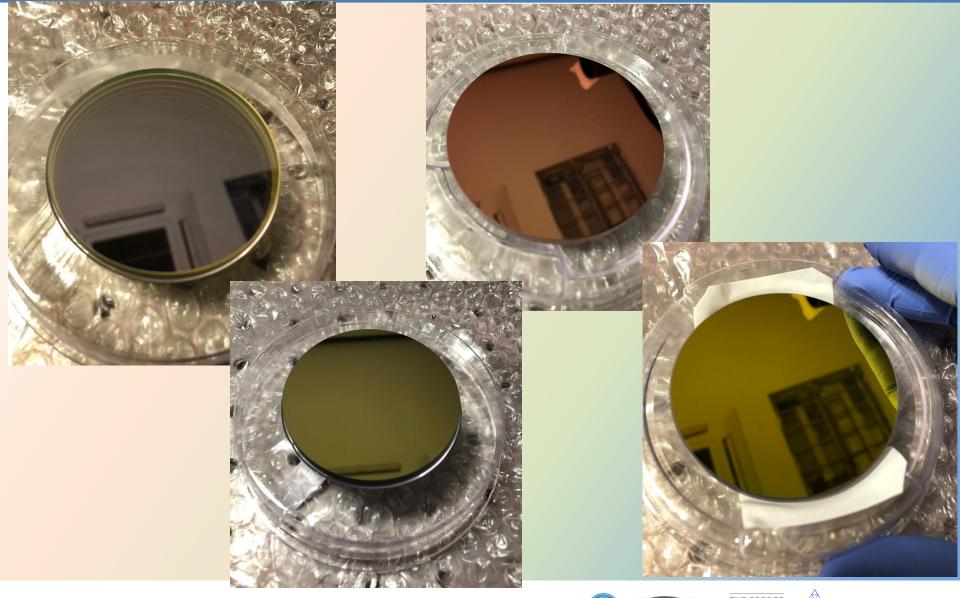






NIR filters: inspection at LICA





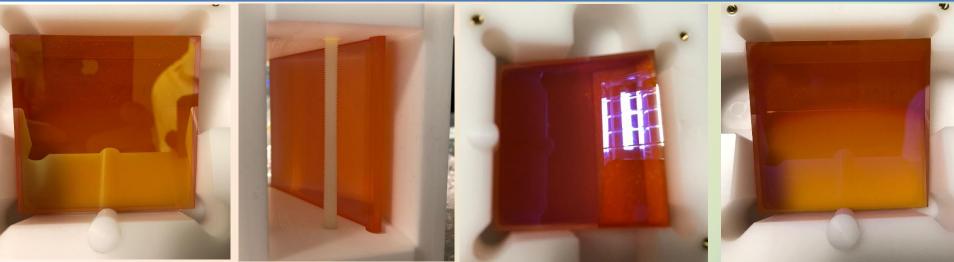






NIR prisms: inspection at LICA

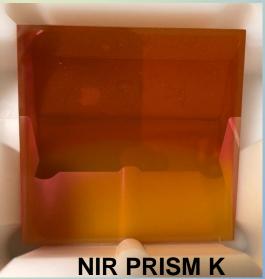




NIR PRISM H



NIR PRISM J





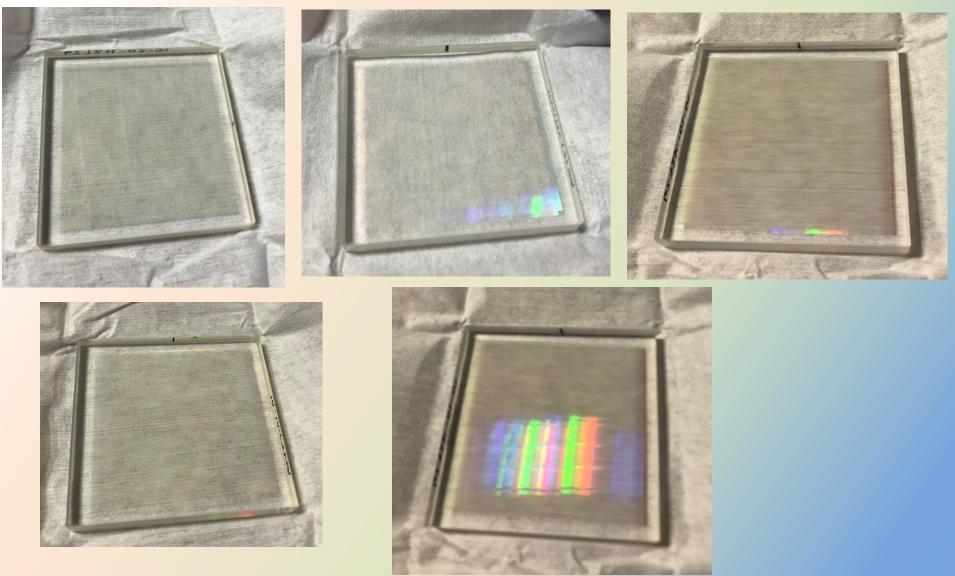






NIR VPHs: inspection at LICA





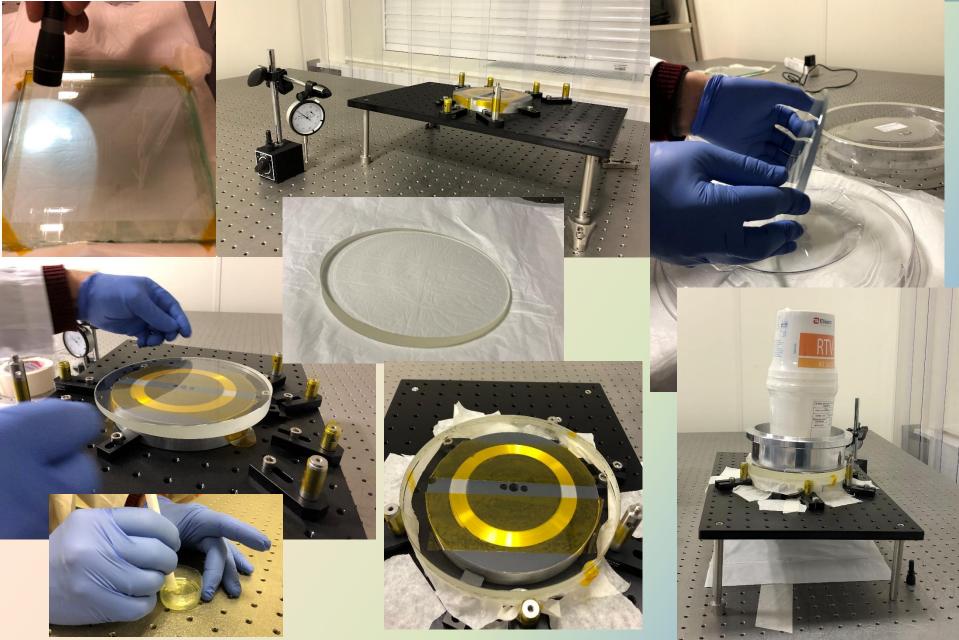






ADC prims gluing







VPH-g bonding

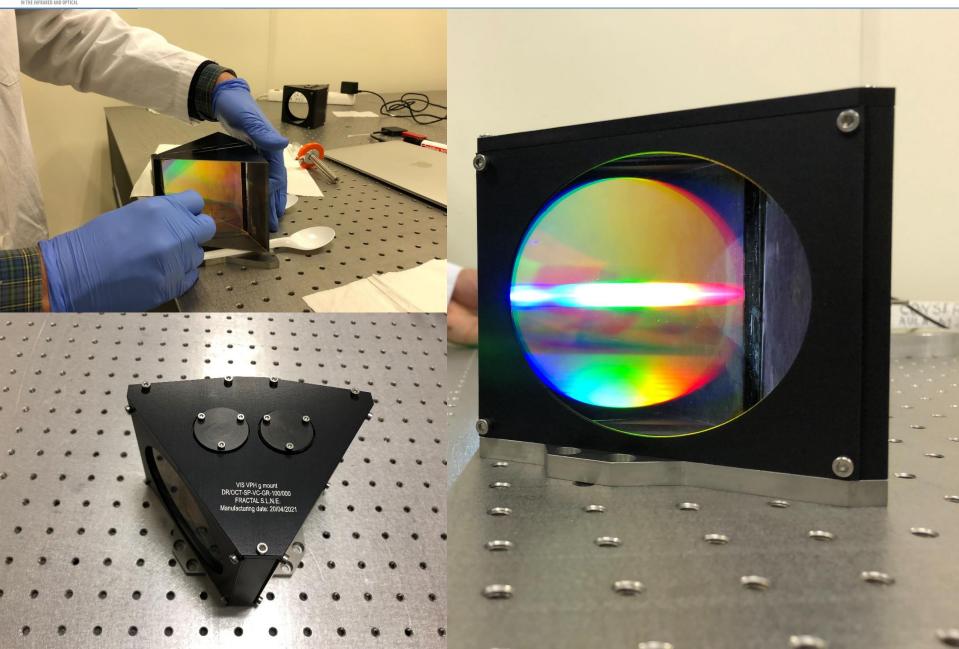






VPH-g Optics/Opto-mechanics assembly

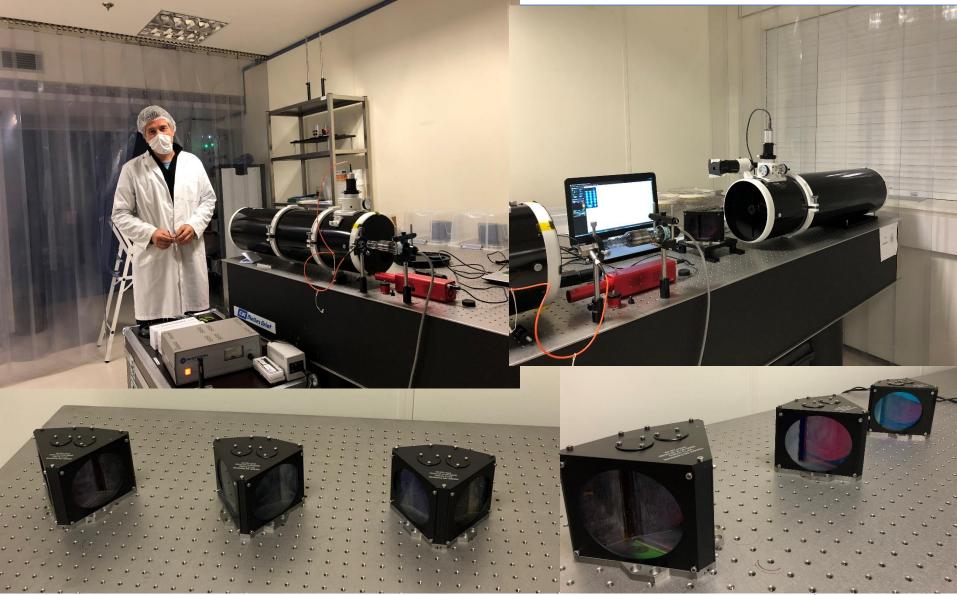






VPH-g, r, i tests





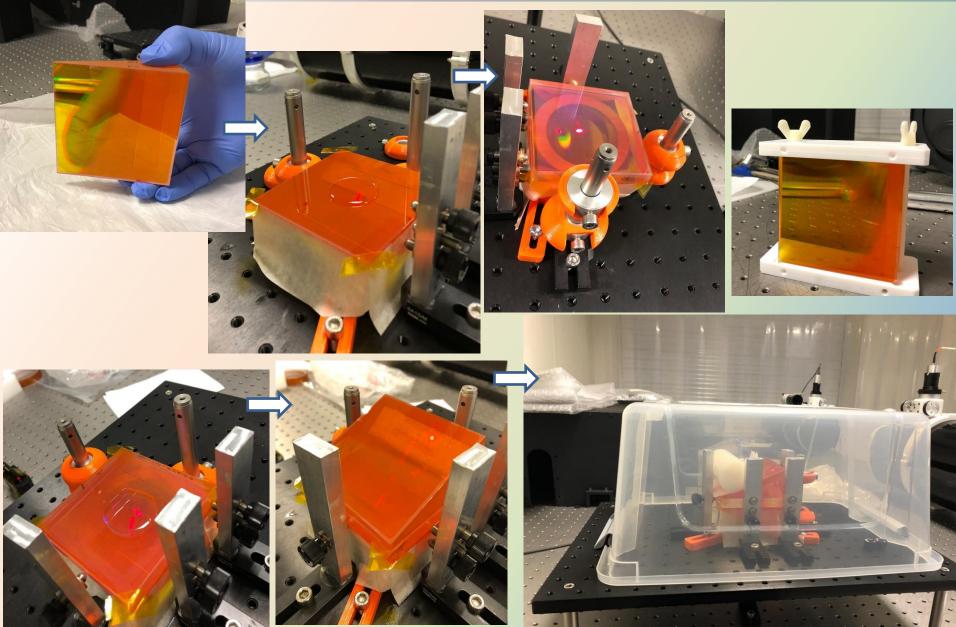






VPH-z bonding

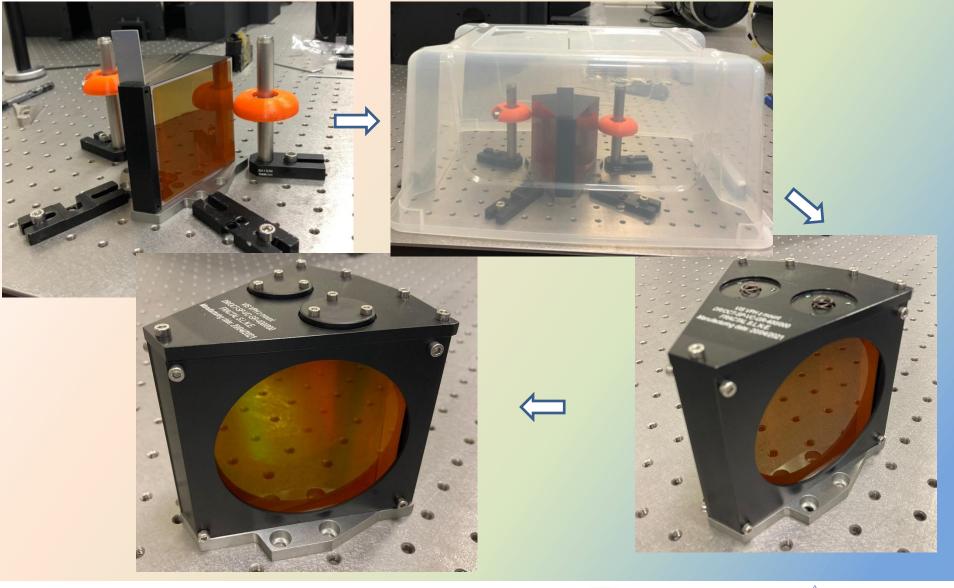






VPH-z Optics/Opto-mechanics assembly





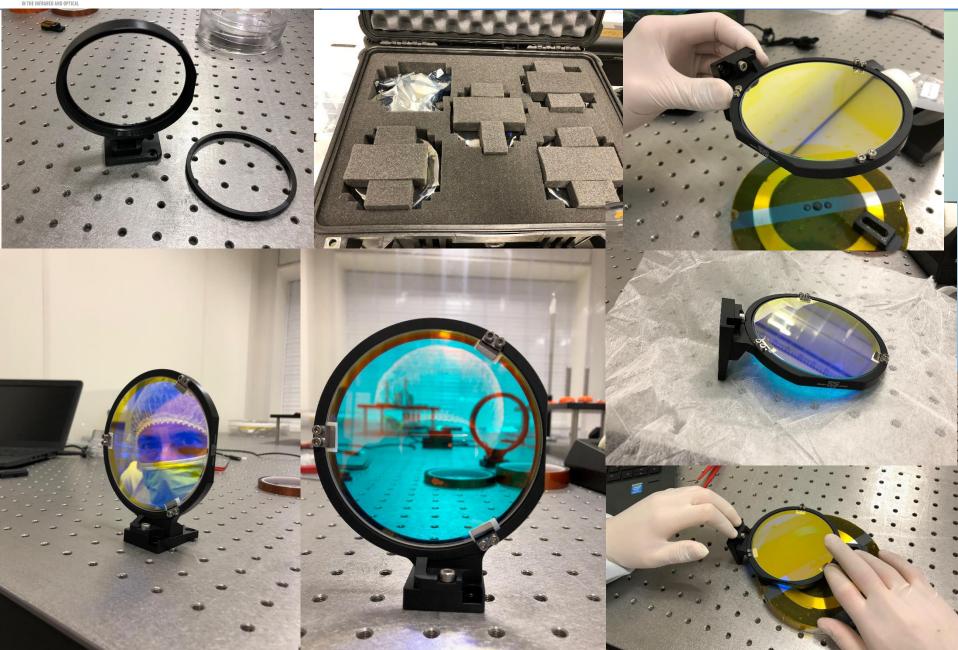




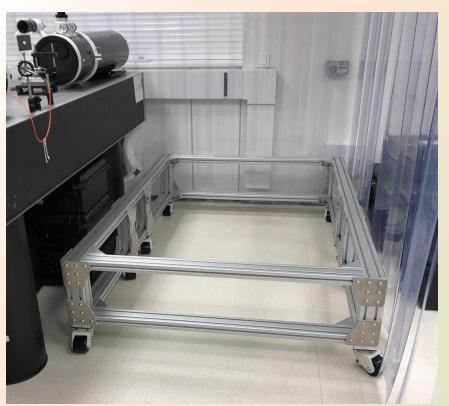


Filter g assembly

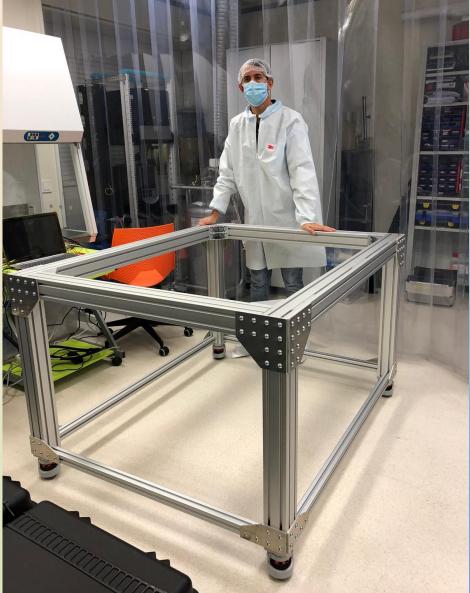




AIV tools. Carts. Finished at LICA

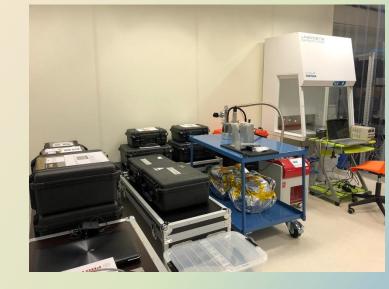


External Frame auxiliary support

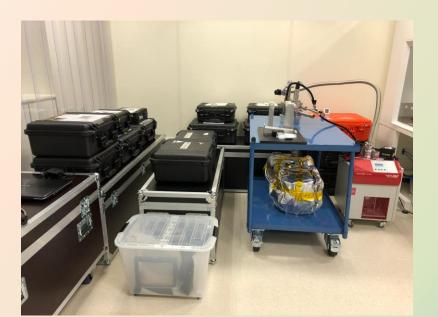


NIR cryostat handling cart





LICA







Anti-Vibration Mounts









Assembled by FRACTAL







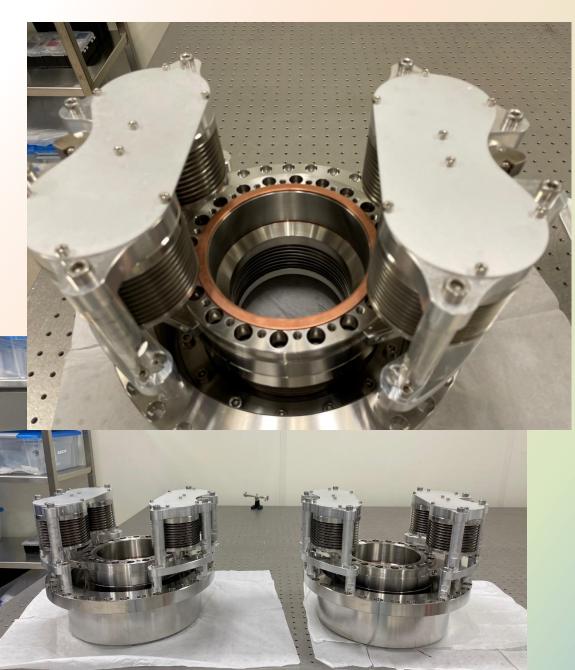








Anti-Vibration Mounts









Getter

Design and assembled by FRACTAL















Air blower and Cover Door custom parts













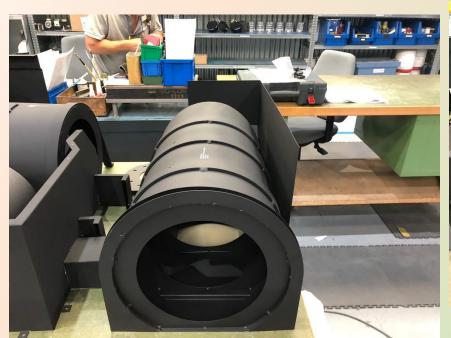
Baffles (at factory, final inspection)









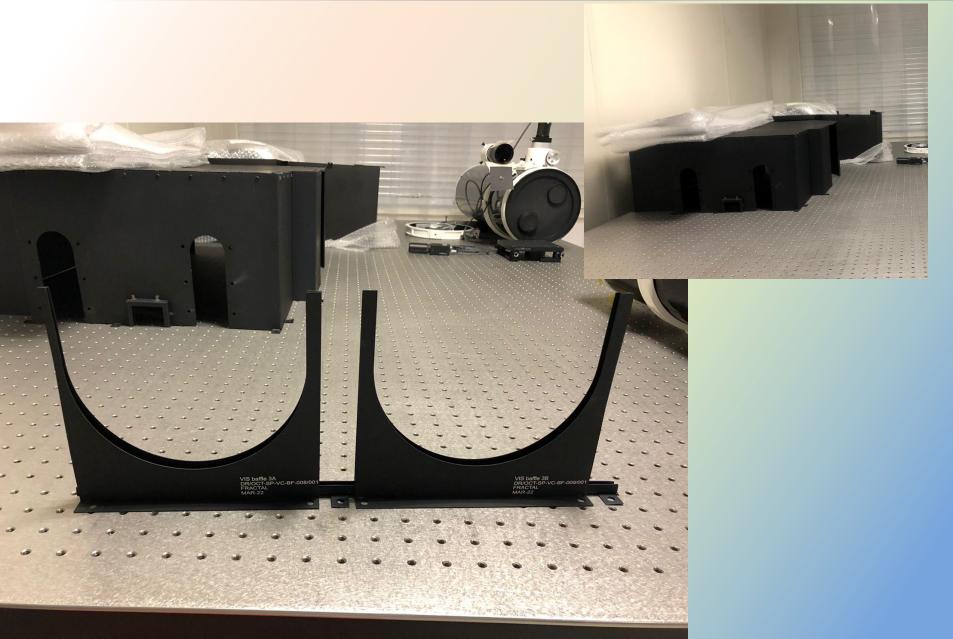






Baffles (at LICA)







Gemini Telescope Simulator

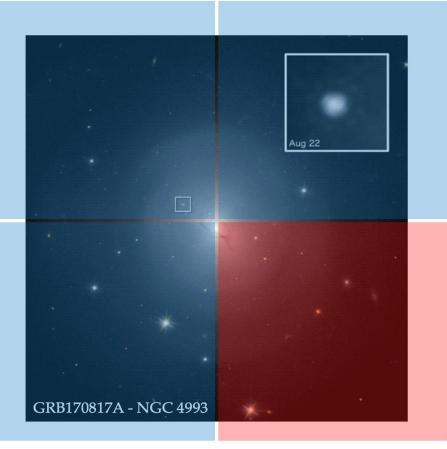






"Tetra-ARmed Super-Ifu Spectrograph"

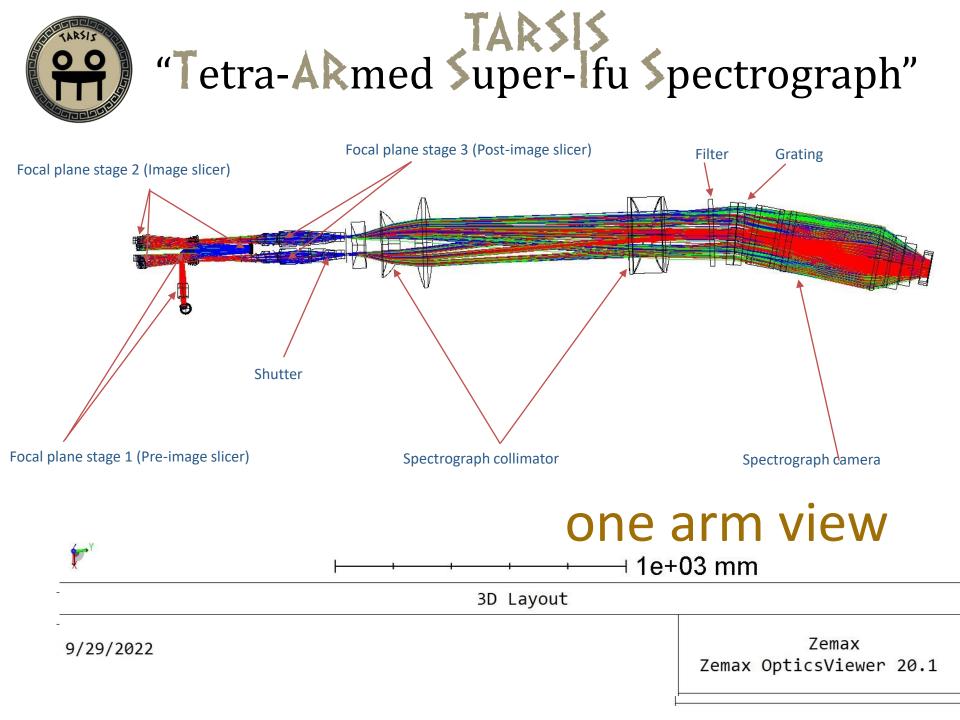
designed for long trips with unknown destinies...

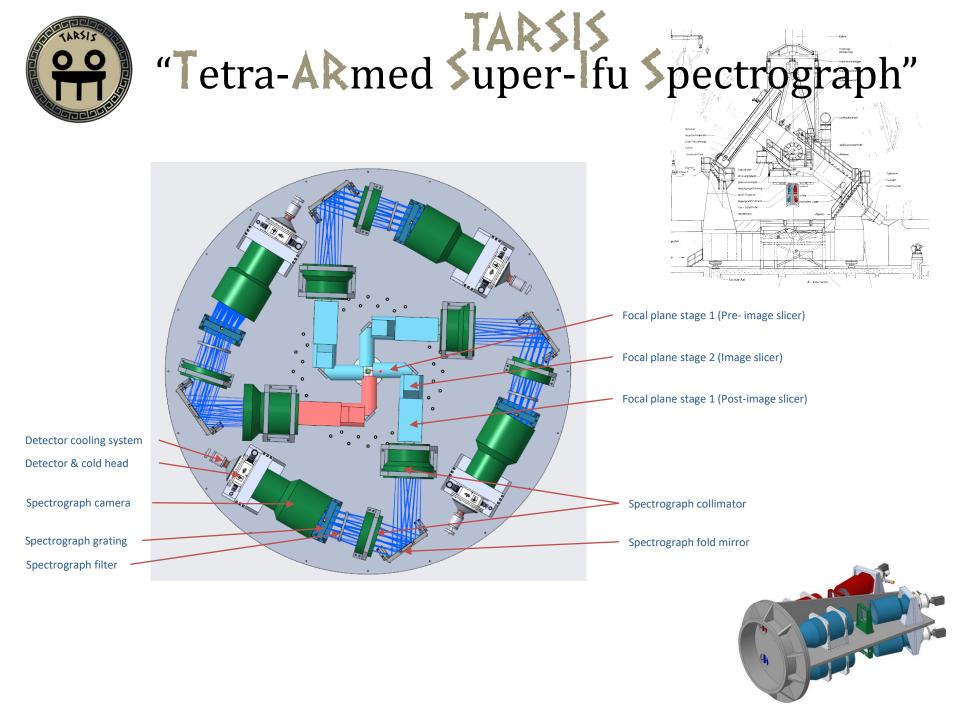


IFU. Connected FoV Large FoV, 9 arcmin² 4 x (1.40 x 1.40) Spaxel size: 2.1 arcsec **Bluest** spectral range Wide λ coverage 320 to 520 nm 510 to 810 nm Spectral resolution R ~1000 **December 2026 + 6 months contingency**

7.2 M€ + 0.70 M€ contingency

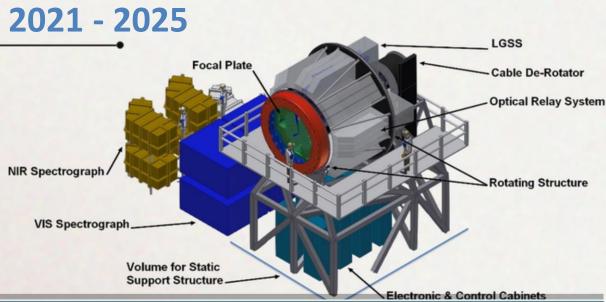






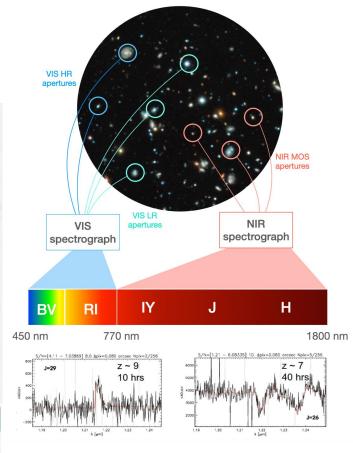


MOSAIC ELT Customer: UCM



nstrument configuration Mode name		Operating wavelength Multiplex (μm)		Aperture on sky (arcsec)	Spectral resolutior (λ/Δλ)
MOS-VIS-LR	MOS	0.45 - 0.7	200	0.7	R ~ 4,000
MOS-VIS-HR	MOS	0.45 – 0.877 100		0.7	R ~ 18,000
MOS-NIR-LR	MOS	0.77 – 1.80 200		0.6	R ~ 4,000
MOS-NIR-HR	MOS	0.77 – 1.80	200	0.6	R ~ 18,000
mIFU-LR	mIFU	0.77 – 1.80 8		2.5	R ~ 4,000
mIFU-HR	mIFU	0.77 – 1.80	8	2.5	R ~ 18,000
Band		Resolution (λ/Δλ)		Bandwidth (µm)	
VIS LR1		4000 (5000 goal)		0.45 - 0.59	
VIS LR2		4 000 (5 000 goal)		0.59 -0.77	
VIS LR3 (Goal)		4 000		0.38-0.45	
VIS HR1		18 000		0.51-0.57	
VIS HR2		18 000		0.61-0.67	
he full near-IR band spanning from 0.77 to 1.	.8µm is covered simultaneously at low spec	stral resolution. The two HR bands in the near-IR	can also be observed simultaneously.		

Band	Resolution (λ/Δλ)	Bandwidth (μm)
NIR LR	4 000 (5 000 goal)	0.77 - 1.8
NIR HR1	9 000	0.77-0.89
NIR HR2	18 000	1.53-1.62



WPs

CALEMOS, Calibration module

NIR spectrographs mechanisms and VPH gratings

CTA Maintenance Plan



Grouped task code	Detailed preventive task code	Grouped task description	Components affected and location/access tools	Development of a Maintenance plan for the, Camera, Bogies and Azimuth Locking system of the LST-1 (Large-sized telescope) of the Cherenkov Telescope	
	Camera regular visual inspection (that needs hardware disassembly) every 6 months from the camera access tower to check:			Array project (CTA)	
CAM-P2	CAM-P-003D	Cooling system fans: regular maintenance (cleaning, checking electrical and data connectors, correct air flow, unforeseen vibrations or noises, etc.)	Cooling system fans (inside camera, front door access)/ need scaffolding	Product Tree FMECA Maintenance	
	CAM-P-004D	Cooling system fans: check power and data connectors, power distribution board and connection to I/O modules	Cooling system fans (inside camera, front door access)/ need scaffolding	Plan	
	External water syste	em: regular maintenance of comme the azimuth rail platfor	Row		
САМ-РЗ	CAM-P-030D	External water system: regular maintenance of commercial chiller as per manufacturer's recommendations. Regular visual inspection to check any leak, pressure and chiller control panel alarms.	Azimuth rail area platform		



Extract from Camera System preventive tasks summary









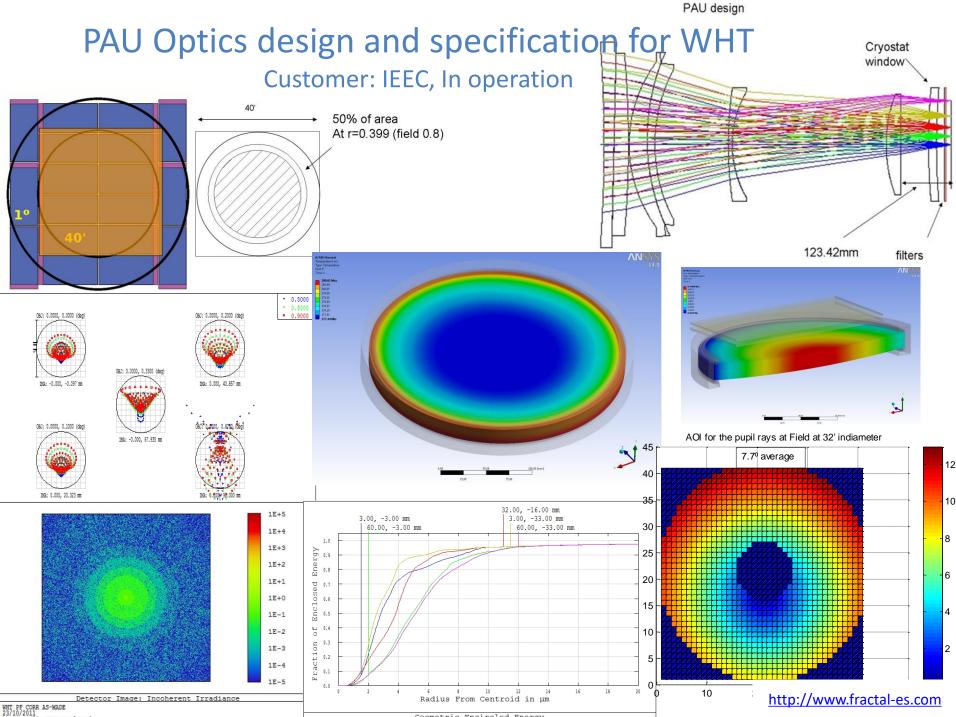




Astronomical instrumentation

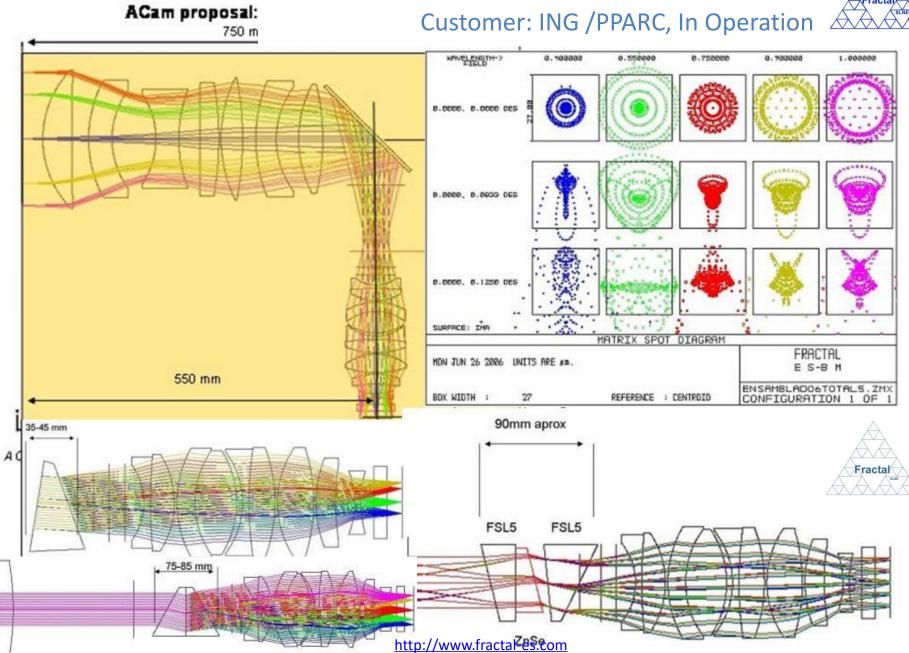
- Design
- Ordering and manufacturing tracking
- Assembly, Integration and Verification

(slides of different projects at the end of the presentation)

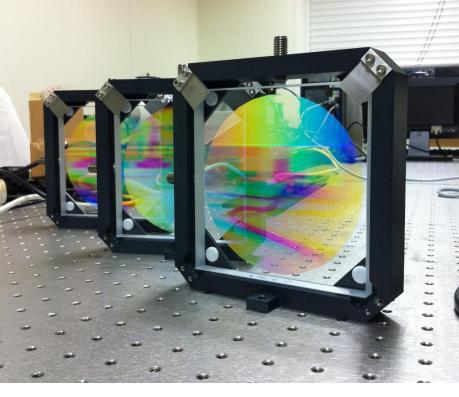


Conceptual Design for ACam (@WHT

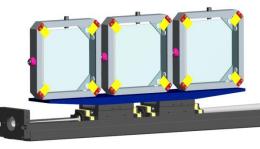


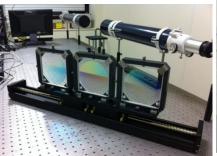


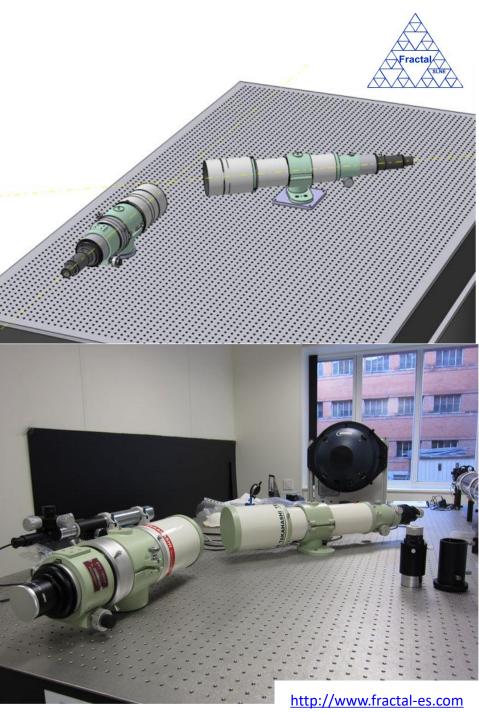




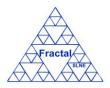
ARES spectrograph Customer: IEEC, in operation

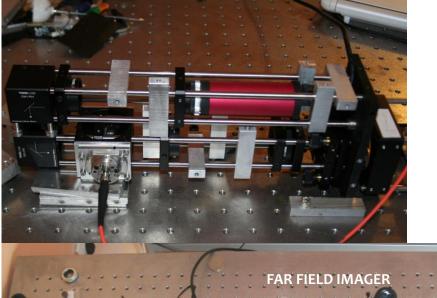




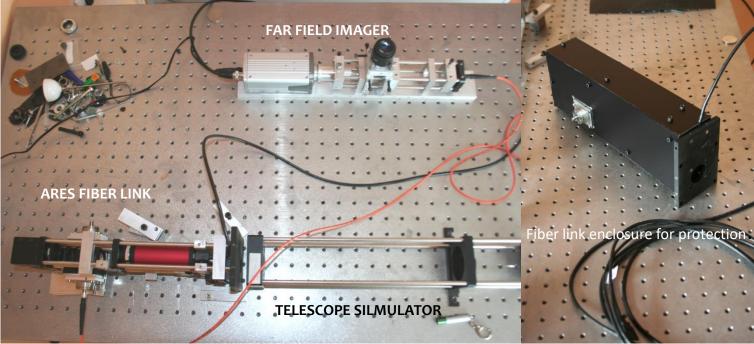


Telescope illuminator + ARES fiber link+ Far field imager. From this setup we obtain the final image after the fiber link

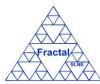


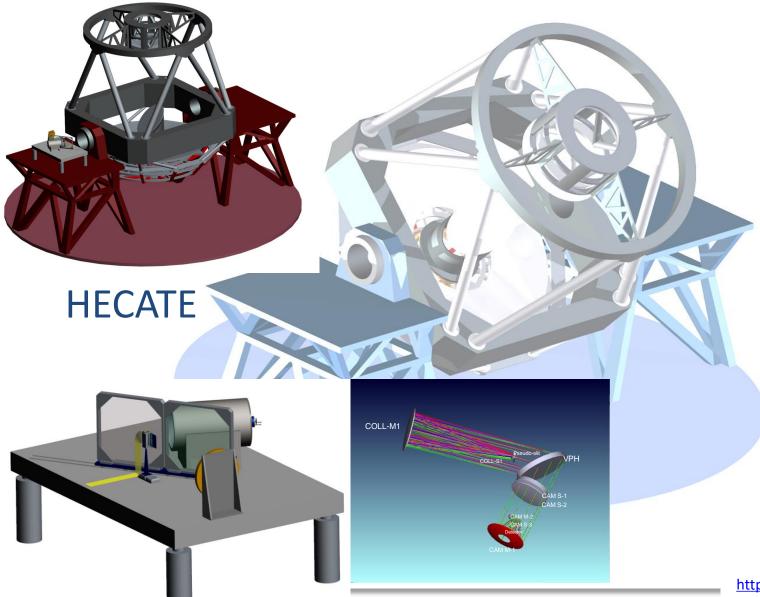


ARES Fiber link Customer: IEEC, in operation



HEXA 6.5m Telescope Conceptual Design Customer: Calar Alto Observatory

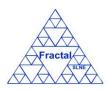


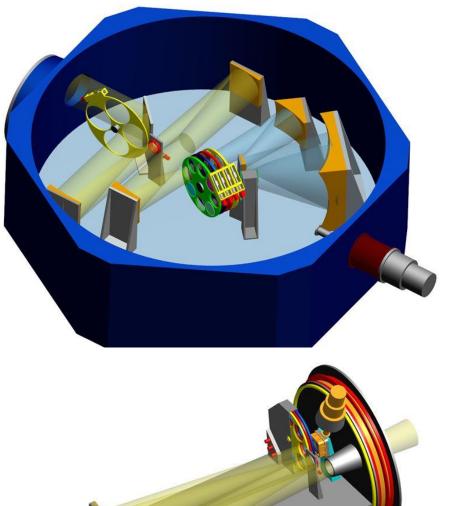


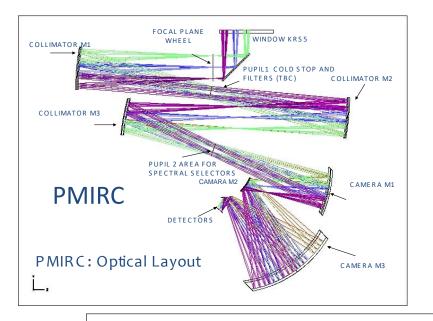
http://www.fractal-es.com

Conceptual Design of IR Instrumentation for Antarctica Customer: Universidad Autónoma de Madrid

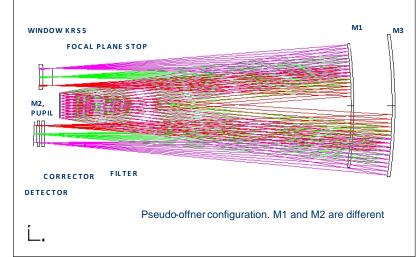
http://www.fractal-es.com

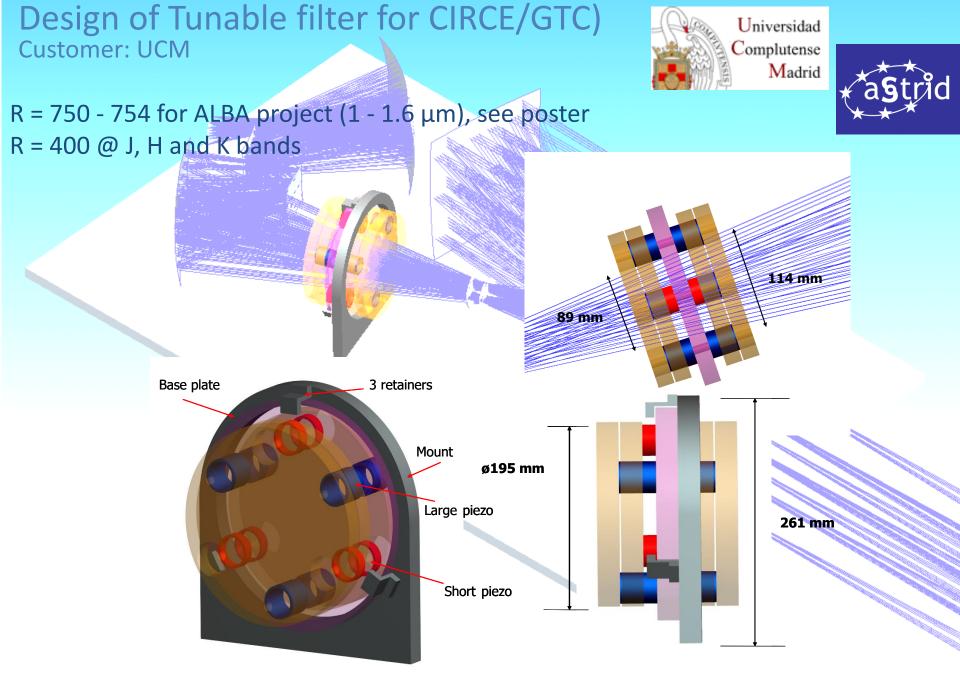






PNIRC





http://www.fractal-es.com





Optics and Opto-mechanics

Optics and Opto-mechanics

Design, specification, manufacturing and tests of optical systems

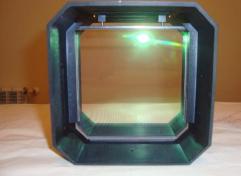
We can carry out the development of any optical system. Starting from the basic scientific requirements, we can produce feasibility studies, designs (at different levels) or the development of the whole project. We have a network of qualified manufacturers for all elements.

 Specification, design, acquisition, and tests of filters, prisms, VPHs (especially for Astronomy applications).

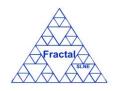


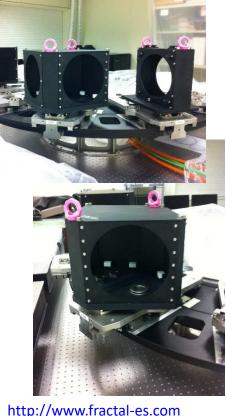


Partnership with Wasatch Photonics

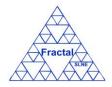


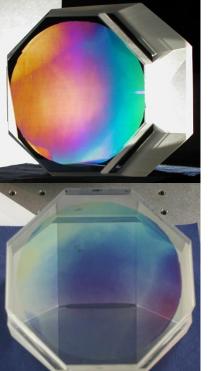




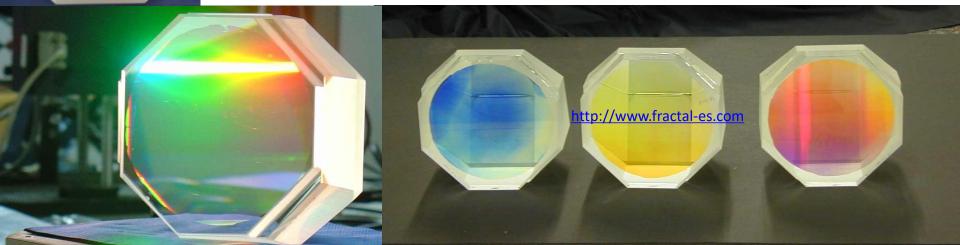


Wasatch Photonics Ultimate Diffraction Gratings

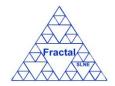




- Wasatch Photonics specializes in high performance Volume Phase Holographic Gratings (VPHGs) and Volume Phase Holographic Optical Elements (VHOEs). Their flagship products are called <u>Dickson® grating</u>s. These are an enhanced form of VPHGs using Wasatch Photonics' patented technologies and proprietary knowhow. <u>http://www.wasatchphotonics.com/index.html</u>
- FRACTAL is partner of Wasatch Photonics for Product Distribution and future R+D. We can specify what our customer need and work with our partner to provide the required gratings
- FRACTAL is distributor for Wasatch Photonics in several countries







VPHs for ExoMars

Diffraction Gratings Specification for RLS Raman Spectrometer

Customer: INTA

The spectrometer unit belongs to the RLS system of the ExoMars-Pasteur rover's payload instruments. The RLS system is part of the Analytic Lab instrument placed inside the rover and provides Raman spectroscopy.

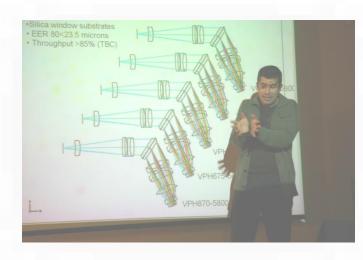




Specialized Training



Training at customer's offices



General 1, 2, 3 days-courses in Madrid

Three perspectives on Project Management

System Engineering: Application and Development

Optical systems design for Instrumentation

Adaptive Optics

Opto-mechanics systems projects

Introduction to Mechanical Design for IR Instrumentation

Optical Fibres in Astronomy

Scientific Imaging in the visible and near-IR

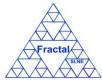
RAMS analyses

Image Processing Techniques

Object Oriented Software Project engineering

CAN/CANOpen buses integration

http://www.fractalslne.es/calendario.jsp





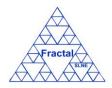


Technology Transfer and Innovation

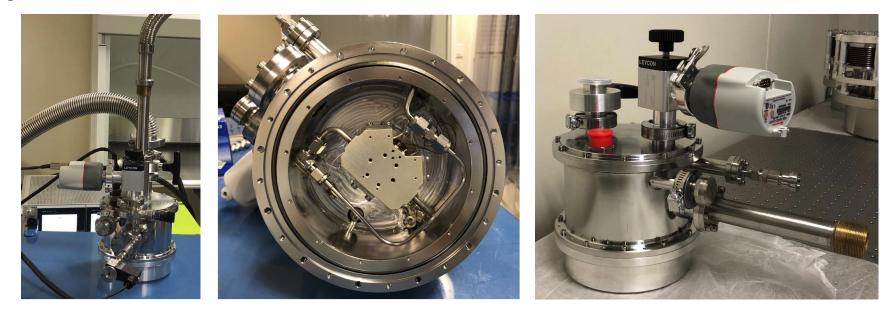
ESO Technology Transfer

(slides of I+D projects at the end of the presentation)

Continuous Flow Cryostats



Technology Transfer Licence Agreement for Cooling System from ESO Signature: 2015



Agreement to license its cooling system technology to our company <u>FRACTAL SLNE</u>. Continuous flow cooling systems were first developed at ESO for the cooling of instruments and detectors for the Very Large Telescope (VLT).

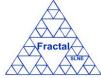
Continuous flow systems are more compact than traditional systems that use a bath of very cold liquid, but share the advantage of being vibration free. The continuous flow system allows the temperature within an instrument to be changed from room temperature to -193 degrees Celsius and is not affected by the loss of electrical power. As a result, the instrument can be kept very close to its operating temperature and the vacuum within the instrument can be retained.

This agreement includes the licensing of the well-proven cryostats (which regulate the low temperatures) that are used to cool individual detectors from room temperature down to -193 degrees Celsius and the more sophisticated version that is used for the cooling of the ultrastable detectors used in more sensitive measurements.

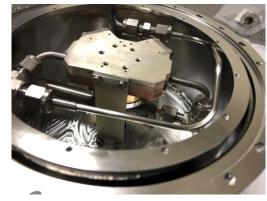


CFCs for ESO

SCANLAF FORTUNA



4 Continuous Flow Cryostats (CFCs) for ESO have been manufactured, integrated and verified.

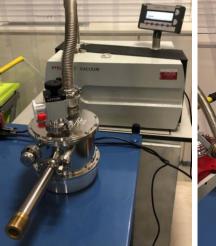








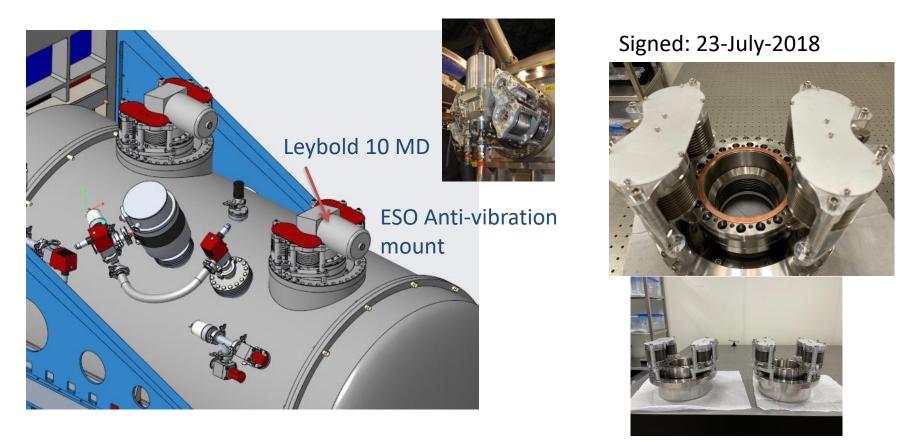








ESO Anti-vibration for Closed Cycle Cooler Cold Head technology



ESO signed an agreement to license anti-vibration mount for Closed Cycle Cooler Cold Head to our company <u>FRACTAL S.L.N.E</u>. The photo on the right (Credit: ESO) shows CRIRES instrument on the VLT. The instrument helium cold heads (Leybold 10 MD) are equipped with anti-vibration mounts to reduce the cold head vibrations to the stability levels demanded for the instrument. The new instrument for Gemini (SCORPIO, shown on the left and currently in detailed design phase) shall include this technology.

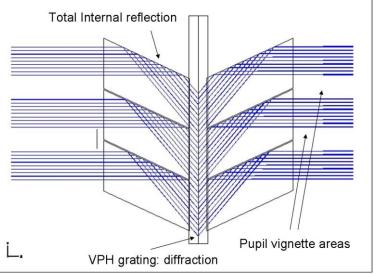
ESO initially developed a damping system that was applied to ISAAC and NACO (which were using the cold head RG 5/80 from Leybold). Successive improvements were performed to cope with the high stability demanded by the VLTI configuration. Antivibration mounts were also used for HAWK-I and KMOS. The current damping system is a 3rd generation design adapted to the Leybold 10 MD cold head and currently in use at CRIRES. This system keeps the cold head fully floating between two soft bellows removing vibration coming from the motion of the heavy displacer.

Innovative Prototypes

Sliced Pupil Grating Prototype



Sliced Pupil Grating Prototype, FRACTAL-UCM





Dirección General de Economía, Estadística e Innovación Tecnológica CONSEJERÍA DE ECONOMÍA Y HACIENDA **Comunidad de Madrid**

Unión Europea





Fondo Europeo De Desarrollo Regional Motivation: to look for a high resolution element when

Goal: to multiply the spectral resolution of your already-built instrument by a factor **of 2 or 3**

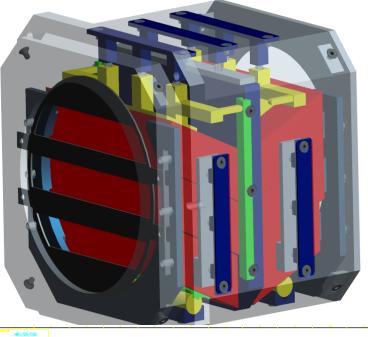
- Your instrument is in Littrow (Ground-based)
- You have problems of weight (Space)
- You have problems of feasibility due to size and mass (Large/Very Large Ground Telescopes)

The solution uses sliced-pupil gratings. There is some vignetting (less than in a non-VPH grating) but can be lowered by increasing the number of slices and the complexity of the integration if needed



This project has been co-funded by CAM [22/2009 for Aero-spatial Innovation, CAM+UE] (51%), FRACTAL (36%) and UCM (15%)





Opto-Mechanics Sliced Pupil Grating Prototype FRACTAL-UCM



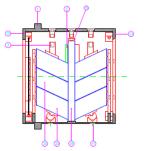
Comunidad de Madrid

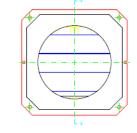


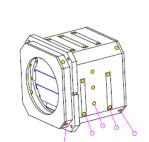


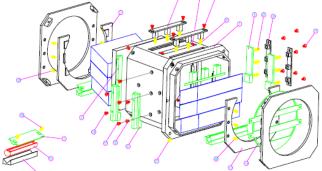
Unión Europea Fondo Europeo De Desarrollo Regional







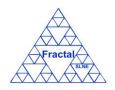










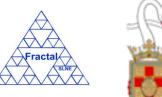


VIENTOS Project

- Partially funded by CDTI under Industry for Science
- The goal is to study novel solutions for pupil elements to be applied to large telescopes (10-m type and E-ELTs)
- Some of the studies within the program are
 - Application of sliced-pupil gratings
 - Tunable filter in the near-IR
 - Filters based on nanotechnology



Centro para el Desarrollo Tecnológico Industrial









Laboratory: Assembly, Integration and Tests

LICA laboratory at UCM Collaborative project FRACTAL – UCM since 2010



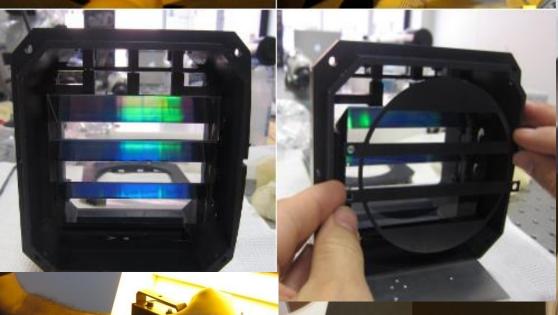




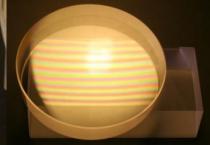


- Instrumentation assembly and tests
- VPH gratings test bench
- Optical bundles test bench
- Detector Test Bench characterization
- Cryogenic system (cryotiger)
- R+D, Innovation and Prototypes
- GTC Control System simulator
- Teaching and practices activities
- Access to UCM facilities

Optical Tests and Assembly



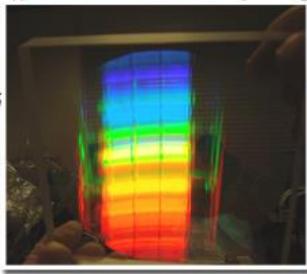


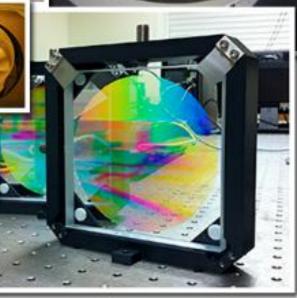












VPH Characterization @ LICA

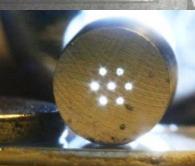




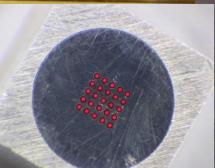


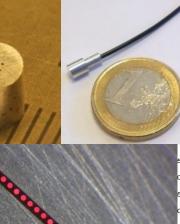


Fiber Tests and bundle assembly

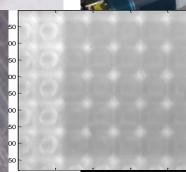


ALTER





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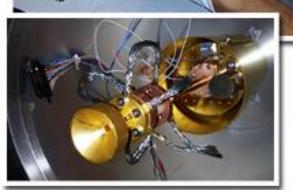




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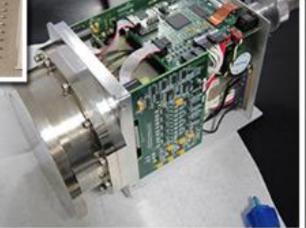




etector Test Bench characterization









Warm test camera





secondary output (fibre optic light guide??)

ND filter wheel controller

Monochromator

Detector Test Bench at LICA-UCM

CCD Characterization

Cold test camera

ARC controller

TEL

0

3

Assembly and tests MEGARA, now SCORPIO starting, then TARSIS and MOSAIC subsystems

1:1

FRACTAL partners have built a small observatory in Sierra de Guadarrama (Madrid) where instrumentation and software can be tested on sky.





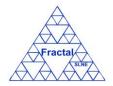
Proprietary Amateur Observatory (Madrid)





Software and Control systems

Software services



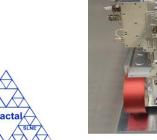
- Real Time Software: Telescopes & Instruments CS
- Final-user S/W for Instruments (mission planning, exposure time calculators, preparation of observations with multi-object spectrographs, data visualization, etc.)
- Databases
- Customized S/W
- In-situ S/W engineering services
- Proprietary S/W for Project Management & System Engineering
- Proprietary Operation and Maintenance Tool

Software development, scientific support and operations on-site services Customers: ESO, OAN, CAB, UCM, IAA

Software Development

- Databases
- Observation planning tools
- Data reduction
- Intranet facilities
- Telescope and instruments control system

Astronomical support for the Spanish National Observatory (Radioastronomy) in the development of a user support platform for the HIFI instrument (Herschel mission, ESA)	CNIG OAN	2006-2010
Pre-launch Operational support for the Spanish National Observatory (Radioastronomy) for the HIFI instrument (Herschel mission, ESA)	CNIG OAN	2006-2010
Software development. Pipelines for VLT instruments (HAWK-I, SINFONI). ESO-Reflex workflows	ESO	2008-2010
Software development. Implementation of data reduction procedures for the ESO VLT pipelines (ANSI-C with the ESO-CPL)	ESO	2010-2012
Proposal for services in ESAC Frame Contract with VEGA-Telespazio	ESA	2017 - 2022





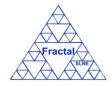
HIFI



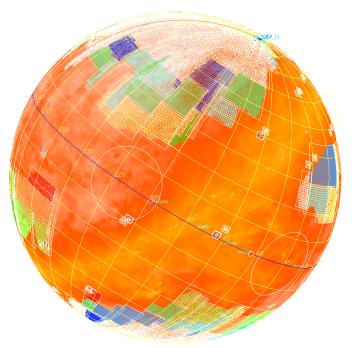


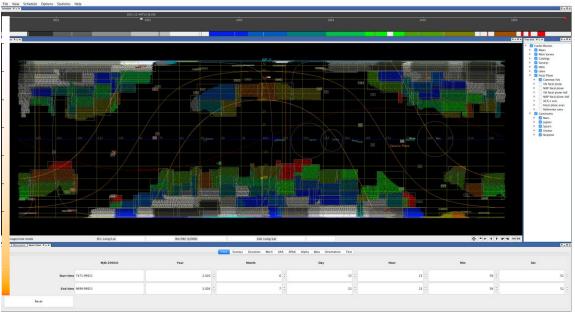


EUCLID Mission Planning Software Customer: Telespazio for ESA



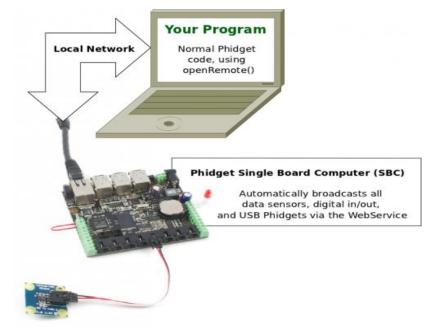
The ESS is the software package in charge of the planning, visualization, validation and inspection of the past, present and future of the mission. It will take the Reference Survey generated from the Euclid Consortium and after validation it will converted to time-tagged pointing commands to ESA's Mission Operation Centre (MOC). The ESS is implemented 100% in java and using GWT for the web components. It is fully OO and implemented using highly reusable components where the basic functionality is available: math, ephemerides, region representation, tessellation (HEALPIX& HTM), DB access, astrometry, etc.





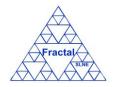
Instrument control system - Technology

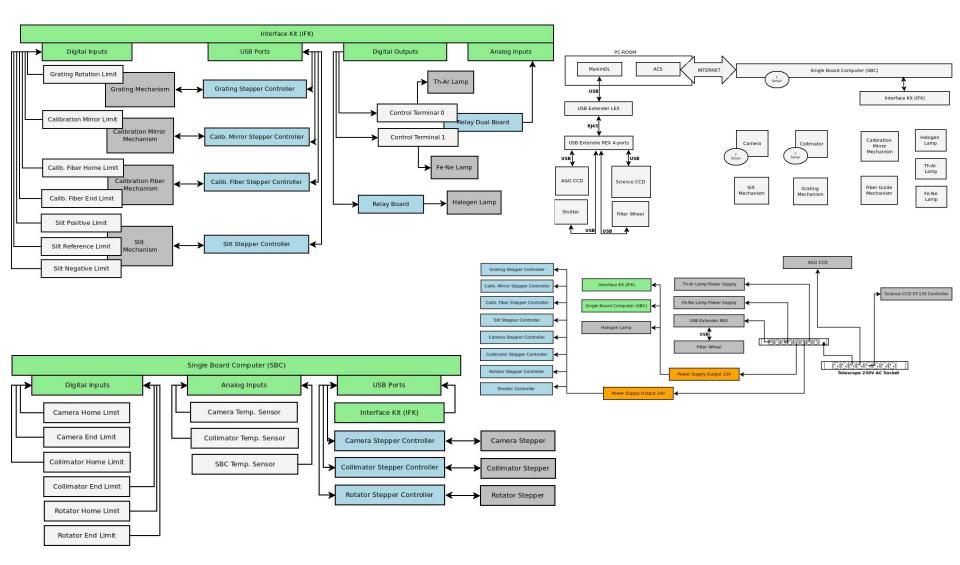
- Phidget products (motor controllers, single board computer, interface kit boards) for controlling all instrument components remotely.
- Can control many motor types:
 - Servo motor
 - Stepper motor
 - DC motor
- Can control many others components:
 - Sensors
 - Relays
 - Limit switches
 - LEDs



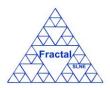
There is no need for proprietary software, Phidgets products work in many popular programing languages (Java, C#, C/C++, Python, Android Java, etc.) and on practically any operating system (Windows, Linux, OS X, Android, iOS, etc..).

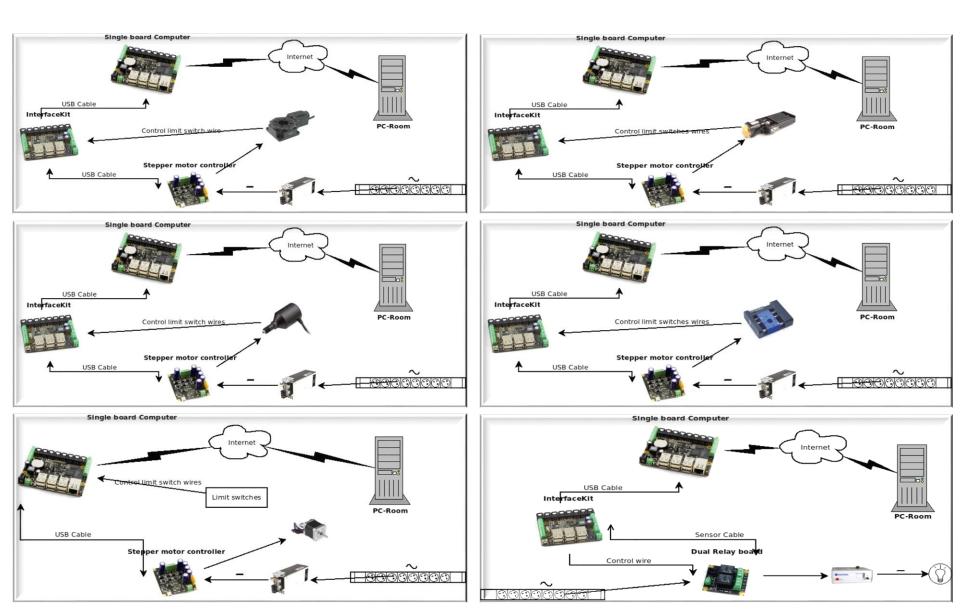
Instrument control system – Architecture



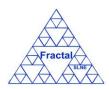


Instrument control system – Hardware



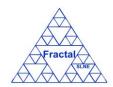


Instrument control system – Software



File Connection	· · ·		File Connection			
Observation Engineering			Observation Engineering			
	Current Status			Current Status		
Observation Mode	Collimator Focusing (mm) Coll. Temp	. (*C) 26.89 ALBIREO Connected	Observation Mode Calibration-Arc Module Calibration	Collimator Focusing (mm)	Coll. Temp. (*C) 26.89	ALBIREO Connected
Slit width (")	Camera Focusing (mm) Cam. Temp	. (°C) 26.22 Grating microsteps	Slit width (*)	Camera Focusing (mm)	Cam. Temp. (°C) 26.22 Grating	microsteps
Grating mounted	Grating wavelength (nm) SBC Temp	. (°C) 30.00 Rotator PA (deg)	Grating mounted 1200b	Grating wavelength (nm)	SBC Temp. (*C) 30.00 Rotat	or PA (deg)
	Observation Setup	Calibration Module	Observat	ion Setup	Calibration	Module
Observer ID	Camera and Collimator focusing	Calibration Mirror	Observer ID	Camera and Collimator fo	ocusing Calibration Mirror	V ON
Observation Mode	Camera 🛆 0.000		Observation Mode Calibration-Arc Module	Camera A	0.000	
	Apply Collimator 0.000	Calibration Fiber Guide ON	Apply	Collimator A	0.000 Calibration Fiber G	uide 🖌 ON
Grating mounted	Paralactic Angle (deg) 0.0 Apply	Tungsten-Halogen lamp ON	Grating mounted 1200b	Paralactic Angle (deg) 0.0 Pply Apply		
CWL (nm) Move to (microsteps)	Apply Apply Sik width (*)	Th-Ar lamp ON		pply Slit width (")	Th-Ar lamp	V ON
Recommended OS filter :	Apply Slit width (*)	Fe-Ne lamp ON	Recommended OS filter : Open	Apply	Fe-Ne lamp	on 📃
Log			Log Wed Sep 21 16:41:47 CEST 2016 INFO: SBC Interfac			
Wed Sep 21 16:41:48 CEST 2016 IN Wed Sep 21 16:41:51 CEST 2016 IN Wed Sep 21 16:42:05 CEST 2016 IN	FG: COLUMATOR stepper controller is attached. FG: GRATING Stepper controller is attached. FG: CALIBRATION_MIRROR stepper controller is attached. FG: CALIBRATION_MEREG.GUIDE stepper controller is attached. FG: SLIT stepper controller is attached.		Wed Sep 21 16:41;47 CEST 2016 INFO: COLLIMATOR Wed Sep 21 16:41;48 CEST 2016 INFO: GARTING ste Wed Sep 21 16:41;51 CEST 2016 INFO: CAUBRATION Wed Sep 21 16:42:05 CEST 2016 INFO: SLUT stepper	pper controller is attached. _MIRROR stepper controller is attached. _FIBER_GUIDE stepper controller is attache	ed.	V
80			80			
File Connection			File Connection			
Observation Engineering			Observation Engineering			
This pa	Engineering Mode nel is only for engineering purposes. The actions can cause damage.	So it's on your own risk	This papel is only for er	Engineering Mode	use damage. So it's on your own risk	
Engineering Password	Login		Engineering Password	Login		
Select stepper motor	Min Position (microsteps) Max Positio	n (microsteps)	Select stepper motor Min Po	sition (microsteps) None	Max Position (microsteps) None	
Calibration Fiber Guide Calibration Mirror		n (microsteps)	Calibration Fiber Guide Calibration Mirror Actual Po	sition (microsteps) -36800	Target Position (microsteps) -36800	
Camera Collimator	Actual Velocity (microsteps/sec) Velocity Limit (m	icrosteps/sec)	Camera		locity Limit (microsteps/sec) 8000.0	Controller
Grating Rotator		rrent Limit (A)	Grating	(microsteps/sec*2) 1.0E7	Current Limit (A) 0.5	connected
Slit		ind limit switch	Slit	Home limit switch No Pressed	End limit switch	
	Actual current (A)	Stop		Actual current (A) -1.730694E-44	Stop	
	Actual current (A)			Actual current (A)		
	Move to (microsteps) Apply Chang	e Actual Position (microsteps) Apply	May	e to (microsteps)	Change Actual Position (microsteps)	Apply
	Velocity Limit (microsteps/sec) Apply	Change Current Limit (A) Apply		(microsteps/sec) Apply	Change Current Limit (A)	Apply
		contraction (microsteps/sec ²) Apply	Velocity Limit	(microsteps/sec) Apply	Change Current Limit (A)	Apply
	Init stepper motor	Stop stepper motor		Init stepper motor		
	init stepper motor	Stop stepper motor		Init stepper motor		itop stepper motor
Log			Log			
Mon Aug 08 14:54:39 CEST 2016	F0: SBC InterfaceKit is connected to the network.		Mon Aug 08 14:54:39 CEST 2016 INFO: SBC Interface	Kit is connected to the network		
Mon Aug 08 14:54:39 CEST 2016 INF Mon Aug 08 14:54:39 CEST 2016 INF	F0: SBC InterfaceKit is connected.		Mon Aug 08 14:54:39 CEST 2016 INFO: SBC Interface Mon Aug 08 14:54:39 CEST 2016 INFO: SBC Interface Mon Aug 08 14:54:39 CEST 2016 INFO: InterfaceKt is Mon Aug 08 14:54:40 CEST 2016 INFO: CALIBRATION_	Kit is connected. connected.		

Software development for Continuous Flow Cryostats (CfCs) technology



In June 2015, ESO and FRACTAL signed a "Non-exclusive know-how license agreement on the use of the ESO Continuous flow cryostat technology".

The concept of this technology is cooled a CCD detector with a nitrogen cryostat which did not need daily refilling and being smaller as possible.

The operating concept is represented in Figure 1 where DH are Detector Heads, Tn are temperature sensors, Rn are Resistors, V are valves, Sn are pressure sensors and OPn are overpressure sensors.

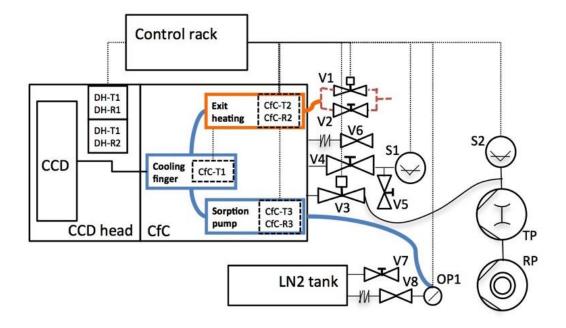
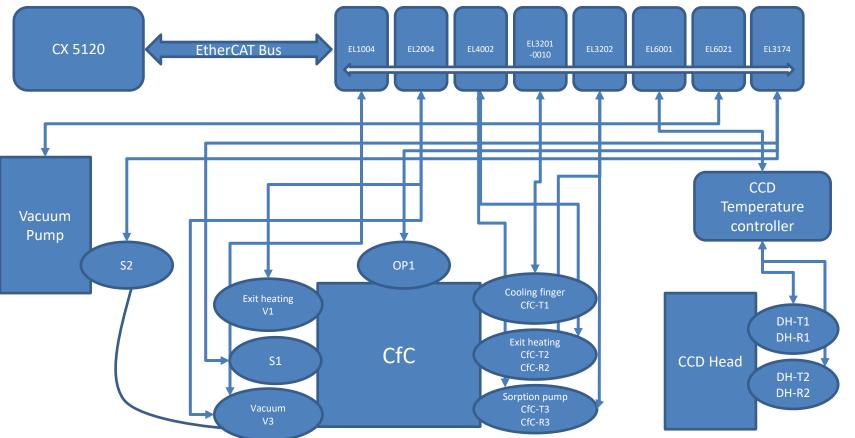


Figure 1: Operation concept of CfCs

CfCs Control System Architecture





	Technology
PLC	Beckhoff CX Embedded PC
Bus	EtherCat
OS	TwinCat 3.1 (Beckhoff)
Programing Language	Structure text (ST), C++
Drivers	N/A (i.e, direct access to HW)



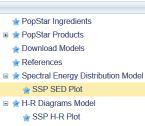
certain units, as described below the plot area. This spectrum can also be plotted normalized to a given wavelength and/or with an offset

RepStar Ingredients

* SSP SED Plot

· The plot can be visualized linear or log-log

A wavelength range for visualization can be defined



- SSP Properties
 - * SSP properties Plot
- Combination of 2 SSP SED Models
 - * Combination of 2 SSP SED Plot
- Photometrical Parameters of a combination of 2 SS
 - reference strategy and the second sec

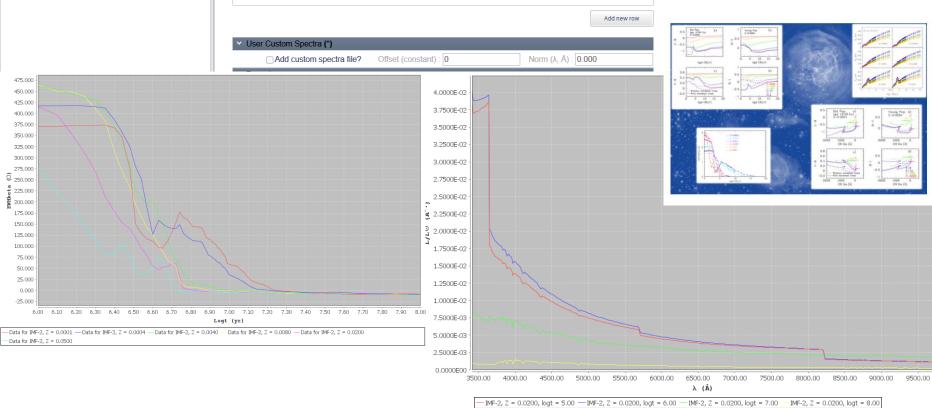
• Once the plot is done, viewing, zooming and printing options will become available by clicking with the right mouse button over the graphic area. Also, it is possible to zoom over a certain plotted area by selecting it with the left mouse button

Eractal Home

ractal

IMF	Z	Log t (yr)	E(B - V)	offset (constant)	Norm (λ, Å)	L type [L/L _☉ , Å ⁻¹]
IMF-2 Salpeter (1955)	0.0200	5.00	0.00	0	0.000	Ltot 🤤
IMF-2 Salpeter (1955)	0.0200	6.00	0	0	0	Ltot 🤤
IMF-2 Salpeter (1955)	0.0200	7.00	0	0	0	Ltot 🤤
IMF-2 Salpeter (1955)	0.0200	8.00	0	0	0	Ltot 🤤

http://www.fractalslne.es/PopStar



Database with the MEGARA-GTC stellar spectral library and a web-based software tool for managing the observations





Sources

Observations

Library Completion

Utilities

Useful links

Project description

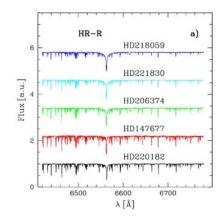
Papers

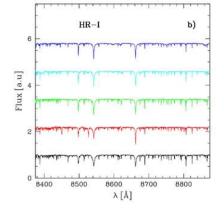
MEGARA-GTC Stellar Spectral Library

MEGARA (Multi Espectrógrafo en GTC de Alta Resolución para Astronomía) is an optical (3650 - 9750Å), fibre-fed, mediumhigh spectral resolution (R = 6000, 12000, 20000) instrument for the GTC 10.4m telescope, in operation since July 2018. The scientific exploitation of MEGARA demands a stellar-spectra library to interpret galaxy data and to estimate the contribution of the stellar populations. To date, there is no complete library, either empirical or theoretical, able to fulfill the MEGARA parameters. The motivation of this project is to obtain a stellar spectral atlas for MEGARA, focused on the highest resolution setups. The spectra have R \sim 20000 in the HR-R and HR-I setups, centred at 6563 and 8633 Å respectively. The data will become available to the community through different releases that can be download here.

The on-going GTC Open-Time program provides the data through filler-type observations in telescope scratch time. The proposals from which we have received granted GTC filler-type time for this program are: GTC22-18B, GTC37-19A, GTC33-19B and GTC19-20A.

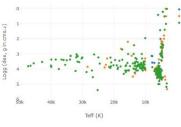
Estimated Date of the release 1.0: June 2020





	7		_
Name			
Other Name			
RA (hh:mm:ss.s) (range between [00:00:11.6,23:59:31.3])			
DEC (±dd:mm:ss.s) (range between [-19:56:35.2,+89:15:50.8])		×	
Δ RA (mas/yr)			
Δ DEC (mas/yr)		-	
SP Type			
Phase-2 (HR-R & HR-I)	Select an Option	*	
Object Type	Select an Option	*	
U (range between [-0.5,16.7])		-	
B (range between [-1.5,19.4])		-	
V (range between [-1.17,18.9])		-	
R (range between [-1.17,18.9])		-	
I (range between [-2.45,16.5])		-	
J (range between [-3,14.1])		-	
Teff (range between [2389,49863])		-	
Teff-L		-	
Logg (range between [-0.25,5.5])			
Logg-L		-	
[Fe/H] (range between [-2.99,2.9])]-[
[Fe/H]-L		-	
Reference			
Comments			
Other Comments			

Name	RA 2000.0 (hhcmmcss.s)	DEC 2000.0 (dd:mm:ss.s)		
** KIR 4	15:23:12.2	+30:17:18.0	modfy	delete
2MASS /15065441+1321060	15:06:54.4	+13:21:06.1	modify	deiete
2MASS /15074769-1627386	15:07:47.7	-16:27:38.6	modify	delete
2MA55 /22244381-0158521	22:24:43.8	-01:58:52.1	modify	delete
80+002058A	07:43:44.0	+00:04:01.0	modify	delete
aD+003740	17:38:37.8	+00:01:44.1	modfy	delete
3D+012916	14:21:45.3	+00:46:59.2	modify	delete
3D+020571	03:37:22.6	+03:16:24.2	modfy	delete
BD+023375	17:39:45.6	+02:25:00.0	modify	delete
3D+024651	23:19:40.5	+03:22:16.7	modify	delete
	2	:41:36.1	modify	delete









Customers: CIEMAT, INAOE



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Sources

Observations

Products

Download

Useful links

Project description

Papers

MEGADES: MEGARA Galaxy Discs Evolution Survey

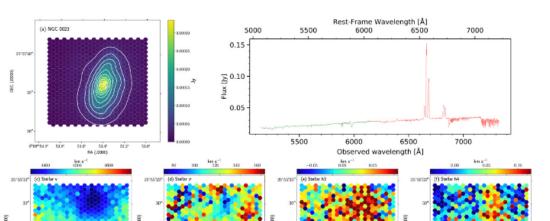
The main interest of the Science Team for the exploitation of the MEGARA (Multi Espectrógrafo en GTC de Alta Resolución para Astronomía) instrument at the 10.4m Gran Telescopio Canarias (GTC) is devoted to the study of nearby galaxies, with focus on the research of the history of star formation, and chemical and kinematical properties of disc systems. We refer to this project as MEGADES: MEGARA Galaxy Discs Evolution Survey.

The initial goal of MEGADES is to provide a detailed study of the inner regions of nearby disc galaxies, both in terms of their spectrophotometric and chemical evolution, and their dynamical characterisation, by disentangling the contribution of insitu and ex-situ processes to the history of star formation and effective chemical enrichment of these regions. In addition, the dynamical analysis of these inner regions naturally includes the identification and characterization of galactic winds (GWs) potentially present in these regions. At a later stage, we will extend this study further out in galactocentric distance.

The first stage of this project encompasses the analysis of the central regions of a total of 43 nearby galaxies observed with MEGARA for ~ 114 hours, including both Guaranteed Time and Open Time observations, using three different VPHs (VPH480-LR, VPH570-LR and VPH675-LR) covering a spectral range from 4350 to 7288 Å at R = 6000. In this webpage we provide a set of all the processed data products available to the community and early results from the analysis of these data regarding stellar continuum, ionized and neutral gas features.

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Data Release DR1 already available here.



Database and Quick Look

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Home Documents Links Wiki Data Archive Data Archive HIFIStars: Late stages in the stellar evolution as seen with Gai: HIFI Guaranteed Time Key Project aimed to gain deeper insight into the structure, thermodynamic and chemistry of circumstellar envelopes and into the mass-loss history of evolved stars. Team: V. Bujarrabal (coordinator), J. Alcolea, J. Cernicharo, L. Decin, C. Dominik, K. Justtanont, A. Marston, G. Melnick, K. Menten, D. Neufeld, H. Olofsson, P. Planesas, M. Pulecka, M. Schmidt, F. Schöle D. Teyssier, R. Waters. Summary: HIFIStars is one of the Guaranteed Time Key Programs proposed by the HIFI construir devoted to observe the nebulae ejected by late type stars, namely AGB, post-AGB, red supergiant, and ye stars. The excelent performances of HIFI, in particular its high spectral resolution, will allow us a very ac the mass-loss processes in these objects. Most stars, including our Sun, will end their iffe ejecting most of the stellar material back to the interstellar which they were formed long time ago. This material forms large shells around the stars, which can be is with HIFI. In the first steps of this mass ejection process, the red glint phase, the circumstellar she specifical and expands at moderate velocities. However, in the very late phases of the stellar fife, the m acquires a stong axial symmetry and much higher velocities. The mechanism responsible for this impres evolution is not yet understood. The precise history of mass-loss, rather than nuclear burning, regulates the ultimate fate of most stars. Thaterial also dominates the gas and dust budget of the galaxy, including its chemical avolution. Hief mass-loss process in detail and thus the late evolution to the interstellar meterial meterstellar meterial back the interstellar meterial and their contribution to the interstellar meterstellar meterial and their controling to the stars.	WBS USB Frequency (GHz) 0.0 HRS1-H USB IF (GHz) 0.0 HRS2-H USB IF (GHz) 0.0 HRS3-H USB IF (GHz) 0.0 HRS4-H USB IF (Ghz) 0.0 HRS4-V USB IF (GHz) 0.0 HRS4-V USB IF (GHz) 0.0 HRS4-V USB IF (GHz) 0.0
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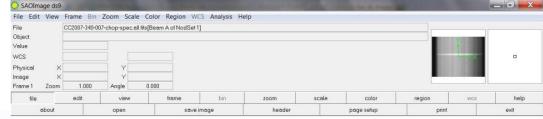
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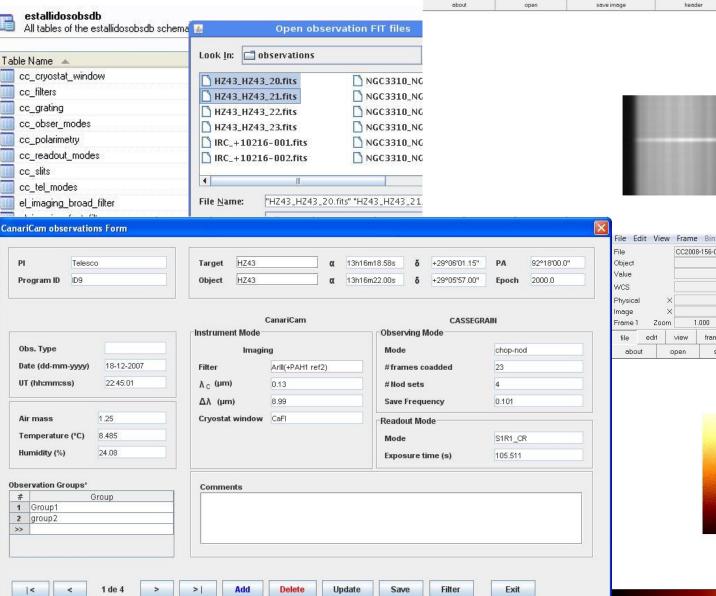
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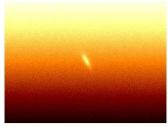
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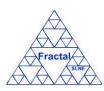


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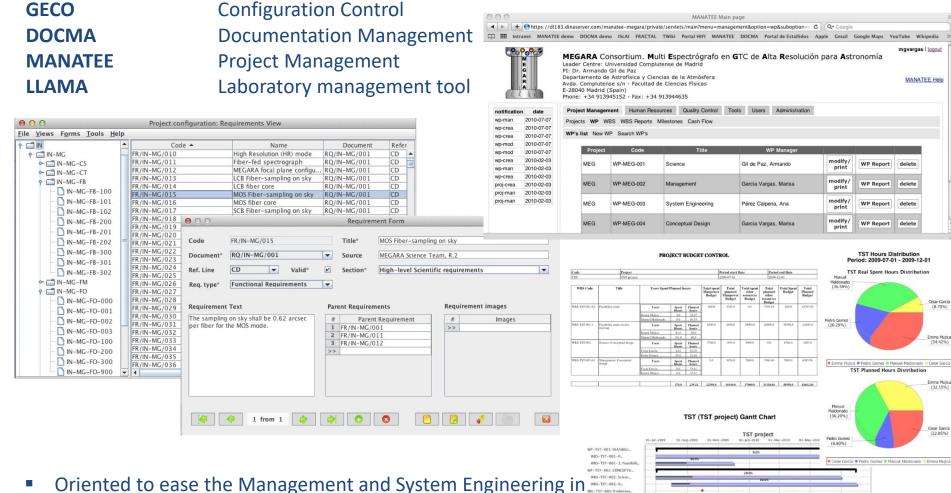
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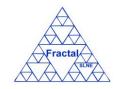


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- Instrumentation projects
 Especially suited for those Consortia and teams collaborating in a multi
 - discipline, complex project in a geographically distributed environment

SUMO: Operation and Maintenance Management tool



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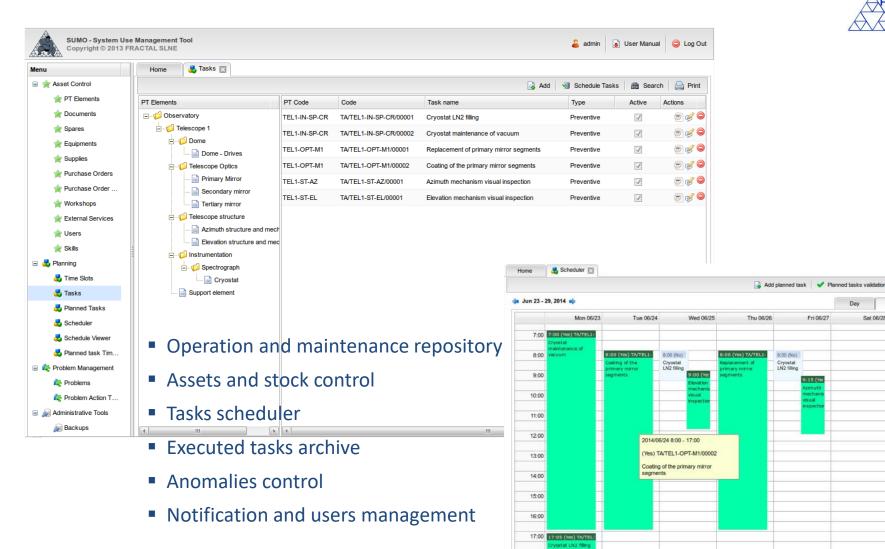
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- Web tool to be useful for distributed organizations.
- Manage the operation and maintenance activities and resources that will be required for the exploitation of a complex system during the Operation phase.

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